

## Study of alpha decay of Superheavy Elements

Shagun Thakur<sup>1\*</sup>, Sushil Kumar<sup>2</sup> and Rajesh kumar<sup>1</sup>,

<sup>1</sup>Department of Physics N.I.T.-Hamirpur, Hamirpur 177005, H.P., INDIA

<sup>2</sup>Department of Physics, Chitkara University Solan 174103, H.P., INDIA

\*email:shagunnitham@gmail.com

### Introduction

Currently nuclear physics is passing through an exciting period of time when one after another new super heavy elements (SHE) are being discovered. These super heavy elements were supposed to undergo spontaneous fission but instead of undergoing fission, each of these elements emits a characteristic  $\alpha$  decay chain, a decay process that was first introduced by Gamow and by Condon and Gurney in the 1920s as a quantum tunneling effect. So, by identifying a particular  $\alpha$  decay chain one can identify a particular super heavy element. The observed decay mode is mainly the  $\alpha$  emission. Alpha Decay is the most prominent tool to investigate the super heavy nuclei [1]. In experiments one usually measures the decay energies and decay times, while the goal of theory is to predict the decay properties of SHE. In 1965 Myers and Swiatecki added the shell corrections to liquid drop model and in 1969 Nilsson et al. predicted that the longest fission half-life centre rather symmetrical around the nucleon numbers  $Z=114$ ,  $N=184$  [2-5]. Such theoretical predictions of long lived super heavy nuclei prompted for a world wide search for such nuclei in nature. Theoretical calculations were performed to predict the absolute alpha decay width, to extract nuclear structure information, and to pursue a microscopic understanding of the alpha decay phenomenon. Recent microscopic calculations predicts long-lived super heavy elements in a variety of shapes, including spherical, axial and triaxial configurations. Many authors proposed that in super heavy mass region the proton numbers 114,120,126 and the neutrons numbers 162,172,184 are the shell closures. In this work, by using Preformed Cluster Model (PCM) the half lives of super heavy elements have been determined and compared with the other existing theoretical results to test the extent of validity of the formalism.

### The preformed cluster decay model

The Half –lifetime according to PCM is given by

$$T_{1/2} = \ln 2 / \lambda$$

Where,

$$\lambda = P_0 v_0 P \quad (1)$$

Here  $v_0$  is the impinging frequency with which the cluster hits the barrier,  $P$  is the penetration probability that gives the probability of penetration of the barrier by the fragment and  $P_0$  is the preformation probability of the cluster giving the probability of the formation of the cluster within the mother nucleus.

For calculating  $P_0$  and  $P$  the dynamical collective coordinate of mass asymmetry  $\eta = (A_1 A_2) / (A_1 + A_2)$  and relative separation  $R$  between the two fragments, is considered via the stationary Schrödinger equation

$$H(\eta, R) \psi_n(\eta, R) = E_n \psi_n(\eta, R) \quad (2)$$

In Principle, the two coordinates are coupled, but in view of the defining equation (1), the Schrödinger equation (2) is solved in the decoupled approximation of  $\eta$  and  $R$ -motions. Only the ground state ( $n=0$ ) solution is relevant for the cluster decay to occur in the ground state of the daughter nucleus. Then, for  $\eta$  motion, the properly normalized fractional cluster preformation probability is

$$P_0(A_2) = |\psi(n)|^2 \sqrt{B_{\eta\eta}(\eta)} / 2A, \quad (3)$$

Details of this model can be found in ref [6]

### Results and discussions

The PCM Based calculations have been made by using the Audi and Wapstra tables of binding energies of 1995 and whenever some values was not present in this table it was taken from the 1995 theoretical binding energies. In table1 we calculate for the alpha decay the half life times for a large number of super heavy elements, though there exist many more [7], which are expected to long enough to detect after the synthesis in the present day experiment set up. We investigate here the magicity (spherical and /or deformed) in superheavy mass region. Any stability against spontaneous fission in this region is due to extra binding resulting from the shell effect which essentially increases the alpha

**Table 1:** TABLE FOR ALPHA DECAY IN VARIOUS ISOTOPES

Z	N	A	P <sub>0</sub>	LOGP <sub>0</sub>	LOGP	Q	Log T <sub>1/2</sub>			
							PCM	VSS	DDM3Y	VS
104	170	274	4.48E-11	-10.3487	-24.924	6.186	13.762	9.21	8.75	9.35
104	174	278	7.89E-12	-11.1029	-30.277	5.216	19.908	14.80	14.31	15.00
104	178	282	4.15E-12	-11.382	-31.853	4.976	21.775	17.88	17.34	18.13
104	182	286	2.74E-12	-11.5622	-33.958	4.686	24.075	23.21	22.65	23.57
104	186	290	2.41E-12	-11.618	-28.725	5.446	18.868	14.67	14.01	14.88
106	172	278	1.92E-10	-9.7167	-22.117	7.006	10.297	7.92	7.49	8.03
106	176	282	9.94E-12	-11.0026	-27.164	5.896	16.67	12.58	12.09	12.74
106	180	286	4.75E-12	-11.3233	-29.103	5.546	18.945	15.61	15.06	15.85
106	184	290	3.1E-12	-11.5086	-29.609	5.456	19.642	18.45	17.87	18.71
106	188	294	2.08E-12	-11.6816	-33.354	4.896	23.586	12.86	12.21	13.03
108	174	282	5.15E-11	-10.2882	-24.213	6.676	12.978	7.13	6.72	7.17
108	178	286	1.97E-11	-10.7055	-23.856	6.746	13.038	8.59	8.11	8.69
108	182	290	5.27E-12	-11.2782	-26.949	6.086	16.728	12.74	12.20	12.92
108	186	294	5.98E-12	-11.2233	-23.502	6.806	13.204	7.35	6.78	7.43
110	176	286	9.98E-11	-10.0009	-21.224	7.606	9.6748	5.38	5.00	5.40
110	180	290	1.39E-11	-10.857	-24.381	6.796	13.715	8.08	7.64	8.11
110	184	294	7.56E-12	-11.1215	-23.685	6.946	13.281	9.67	9.15	9.73
110	188	298	8.22E-12	-11.0851	-20.611	7.736	10.149	6.08	5.54	6.12
112	176	288	4.78E-10	-9.32057	-19.424	8.346	7.1761	2.44	2.14	2.35
112	180	292	3.31E-11	-10.4802	-22.229	7.506	11.165	5.57	5.20	5.56
112	184	296	1.51E-11	-10.821	-21.478	7.696	10.752	6.27	5.83	6.26
112	188	300	1.95E-11	-10.71	-19.323	8.326	8.4711	3.65	3.19	3.59
114	176	290	8.67E-10	-9.06198	-19.58	8.496	7.0707	.02	-.17	-.16
114	180	294	1.70E-10	-9.76955	-19.164	8.616	7.3607	2.84	2.55	2.73
114	184	298	9.01E-11	-10.0453	-18.246	8.916	6.7135	2.98	2.63	2.84
116	168	284	1.44E-06	-5.48416	-13.244	11.606	-2.5569	-6.19	-6.04	-6.57
116	172	288	2.15E-07	-6.66756	-13.661	11.316	-1.3062	-3.18	-3.18	-3.48
116	176	292	7.14E-08	-7.71463	-14.477	10.826	0.000129	-1.99	-2.09	-2.26
116	180	296	1.70E-08	-7.76955	-13.907	11.106	0.049925	-.99	-1.15	-1.25
116	184	300	8.98E-09	-8.04672	-13.333	11.406	-0.25042	-1.02	-1.23	-1.26
118	174	292	1.06E-06	-5.97469	-12.379	12.366	-3.2984	-4.23	-4.15	-4.61
118	178	296	2.74E-07	-6.56225	-12.435	12.286	-2.6515	-3.79	-3.77	-4.15
118	182	300	5.94E-08	-7.22621	-11.792	12.716	-2.6355	-3.56	-3.61	-3.91
120	176	296	1.47E-06	-5.83268	-11.013	13.686	-4.8261	-6.03	-5.84	-6.51
120	180	300	1.61E-06	-5.79317	-11.284	13.396	-4.5876	-5.42	-5.31	-5.87

half lives for nuclei with  $Z < 114$  and  $N < 184$  and decreases those of nuclei with  $Z > 114$  and  $N > 184$ . We Compare here our results with the existed models and found the shell closures at  $Z=114$  and at  $N=184$ . Calculations are given in Table I.

**References**

[1] M. A. Stoyer, Nature 442, 876 (2006).  
 [2] D. N. Poenaru, W.Grenier Phys. Scr 44, 427 (1991)

[3] D. N. Poenaru, W. Grenier J. Phys. G: Nuc. Part. Phys 17, 443 (1991)  
 [4] D. N. Poenaru, W. Grenier, E. Hourani Phys. Rev. C 51, 594 (1995).  
 [5] K. P. Santosh, Antony, Joseph Pramana. J. Phys. 59, 599 (2002).  
 [6] Raj.K Gupta et al J.Phys.G.28,2875-2884(2002)  
 [7] G. Royer, J. Phys. G 26, 1149 (2000).