

## Study of Cluster Emission in Ra Isotopes using Modified Universal Decay Law

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Cluster radioactivity was first predicted by Sandulescu *et al.* in 1980 [1]; these clusters were heavier than  $\alpha$  particles and lighter than fission fragments. However, the experimental proof of this decay was given by Rose and Jones in 1984 [2]:  $^{14}\text{C}$  emitted from  $^{223}\text{Ra}$ . Till now, many clusters decay from light to heavy clusters ( $^{14}\text{C}$  to  $^{32}\text{Si}$ ) have been observed from various trans-lead nuclei (Fr, Ra, etc.) resulting the corresponding daughter as magic ( $Z=82$ ) or neighboring nuclei which indicate the importance of this particular ( $Z>82$ ) region in cluster radioactivity [3].

The half-lives of cluster decay are sensitive to the disintegration energy ( $Q$ -value), which is given by:

$$Q = B.E.(d) + B.E.(c) - B.E.(p) + k[Z_p^\epsilon - Z_d^\epsilon]$$

where  $B.E.(p)$ ,  $B.E.(d)$  and  $B.E.(c)$  are the binding energies of the parent nucleus, daughter nucleus, and emitted cluster, respectively. The term  $k[Z_p^\epsilon - Z_d^\epsilon]$  indicates screening effect caused by the surrounding electrons around the nuclei [4] with  $k=8.7$  eV [ $8.7 \times 10^{-6}\text{MeV}$ ] and  $\epsilon=2.517$  for  $Z \geq 60$ .

Although, cluster decay half-lives can be estimated with the help of various models [5] but there are also various formulas which are capable to calculate the cluster decay half-lives accurately [6]. In 2009, Qi, proposed a formula for the calculation of half lives of  $\alpha$  and

cluster emission. This is also known as universal decay law [8] which was recently modified by Soylu and Qi [9] by adding angular momentum ( $l$ ) and isospin ( $I$ ) dependent terms, named as MUDL and represented as,

$$\log_{10}T_{1/2}(\text{Sec.}) = aZ_cZ_d\sqrt{\frac{\mu}{Q}} + b\sqrt{\mu Z_cZ_d} \\ \sqrt{(A_d^{1/3} + A_c^{1/3}) + c + d} \\ \sqrt{\mu Z_cZ_d(A_d^{1/3} + A_c^{1/3})} \\ \sqrt{l(l+1) + e\sqrt{I(I+1)}}$$

In this formula,  $Z_c$  and  $Z_d$  are the proton number of cluster and daughter nucleus, respectively.  $\mu$  is reduced mass defined as  $A_dA_c/(A_d + A_c)$ . Coefficients of MUDL are given as:  $a = 0.2760038$ ,  $b = -0.2414764$ ,  $c = -42.3417001$ ,  $d = 0.0011798$ , and  $e = 49.66757757$ .

In the present work, we have attempted to find various clusters which have the chances of emission from Ra ( $Z=86$ ) isotopes ( $126 \leq N \leq 147$ ) considering the proton shell closure effect. Therefore, we have calculated cluster decay half-lives for various clusters (isotopes of C:  $Z=6$ ) from Ra isotopes by using MUDL formula given above. For these theoretical estimates, the disintegration energies ( $Q$ -values) are used from WS4 mass model [10], which are found to be more precise than other widely used theories, as shows in our previous work [11]. Consequently, we have found the significant probability of  $^{12-16}\text{C}$  emission from  $^{220-224}\text{Ra}$ , respectively.

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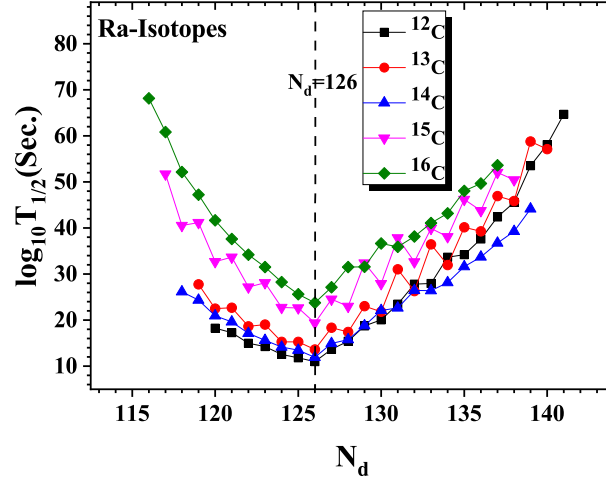


FIG. 1: Variation of half-lives of  $^{12-16}\text{C}$  cluster emissions from Ra isotopes ( $126 \leq N \leq 147$ ) as a function of neutron number of daughter nuclei. These half-lives are calculated by using MUDL formula.

For this study, the logarithmic half-lives of unknown cluster emissions (isotopes of C) from Ra isotopes ( $126 \leq N \leq 147$ ) are plotted in Fig. 1 (up to  $10^{70}$  Sec.) where the minima correspond to  $^{208}\text{Pb}$  daughter i.e., doubly magic ( $Z=82$ ,  $N=126$ ). These minima provide us the

TABLE I: The calculated logarithmic half-lives of highly probable clusters emitted from Ra isotopes ( $126 \leq N \leq 147$ ) using MUDL formula.

Parent nucleus	Daughter nucleus	Emitted cluster	Q (MeV)	$l$	$\log_{10} T_{1/2}$ (Sec.)
$^{220}\text{Ra}$	$^{208}\text{Pb}$	$^{12}\text{C}$	32.13	0	10.98
$^{221}\text{Ra}$	$^{208}\text{Pb}$	$^{13}\text{C}$	31.70	3	13.63
$^{222}\text{Ra}$	$^{208}\text{Pb}$	$^{14}\text{C}$	33.16	0	11.90
$^{223}\text{Ra}$	$^{208}\text{Pb}$	$^{15}\text{C}$	29.22	2	19.38
$^{224}\text{Ra}$	$^{208}\text{Pb}$	$^{16}\text{C}$	26.99	0	23.75

most probable clusters having chances of emission from the respective isotopes. Our main results (most probable clusters) are tabulated in Table I and are within the experimental reach, which are also found in the close match with the recent predictions of Refs. [12, 13].

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