

Sgnificance of nuclear level density and γ -ray strength function in neutron capture $^{68}\text{Zn}(n,\gamma)^{69}\text{Zn}$ reaction cross-section

Enakshi Senapati^{1*}, Satabdi Mondal¹, Srijit Bhattacharya²,
Deepak Pandit^{3,4}, Le Tan Phuc^{5,6}, Nguyen Ngoc Anh⁷, Tran Vu
Dong^{5,6}, Nguyen Dinh Dang⁸, Nguyen Quang Hung^{5,6*}, Balaram Dey¹

¹*Department of Physics, Bankura University, Bankura, West Bengal-722155, India*

²*Department of Physics, Barasat Govt. College, Barasat, N 24 Pgs, Kolkata - 700124, India*

³*Variable Energy Cyclotron Centre, 1/AF-Bidhannagar, Kolkata-700064, India*

⁴*Homi Bhabha National Institute, Training School Complex, Anushaktinagar, Mumbai, 400094, India*

⁵*Institute of Fundamental and Applied Sciences, Duy Tan University, Ho Chi Minh City 70000, Vietnam*

⁶*Faculty of Natural Sciences, Duy Tan University, Danang City 55000, Vietnam*

⁷*Phenikaa Institute for Advanced Study (PIAS), Phenikaa University, Hanoi 12116, Viet Nam*

⁸*Nuclear Many Body Theory Laboratory, RIKEN Nishina Center for Accelerator-Based Science, RIKEN, 2-1 Hirosawa, Wako city, Saitama 351-0198, Japan*

**Email: enakshisenapati2194@gmail.com*

1. Introduction

In nuclear astrophysics, neutron capture reactions in s- and r-processes of nucleosynthesis play a decisive role in the understanding of origin of elements heavier than iron. The description of these neutron capture reactions relies on the statistical Hauser-Feshbach [1] theory to estimate the reaction cross-section. The model requires two indispensable input quantities [2], nuclear level density (NLD) and gamma ray strength function (γ -SF). These parameters however, are rather poorly constrained due to the absence of proper experimental data. So, a good knowledge of NLD and γ -SF is required to get accurate results while using statistical model calculations. The application of strength function is not only important in nucleosynthesis but also in determining soft dipole modes in neutron rich nuclei essential to understand the reaction rate.

2. Methods

Thus at first, calculations are carried out to estimate the $^{68}\text{Zn}(n,\gamma)^{69}\text{Zn}$ reaction cross-section [3] by using the combinations of existing NLDs and γ -SFs in TALYS-1.95 code [4]. It is

seen that most of the combinations either over predict or under predict the experimental data. Therefore, the microscopic EP+IPM [5] (exact pairing plus independent particle model) and EP+PDM [5] (exact pairing plus phonon damping model) have been carried out to calculate the NLD and γ -SF of ^{69}Zn nucleus, respectively. As EP+IPM describes reasonably well the experimental NLD data [6] and EP+PDM calculation gives a good knowledge of γ -SF [7], so these two approaches are applied to TALYS-1.95 code to understand their applicability and reliability in explaining the astrophysical reaction cross-section.

3. Result & discussions

It is seen that microscopic EP+IPM NLD and EP+PDM γ -SF explains the experimental data better than all other combinations available in TALYS-1.95, indicating the impact of the exact treatment of thermal pairing correlation. Furthermore, the inclusion of an UB structure in the EP+PDM γ -SF improves the comparison with the experimental cross-section data in the low energy region 0.01-0.15 MeV, while the calculation without this upbend structure slightly underpredicts the measured data [8]. The result will be further discussed in details at the symposium.

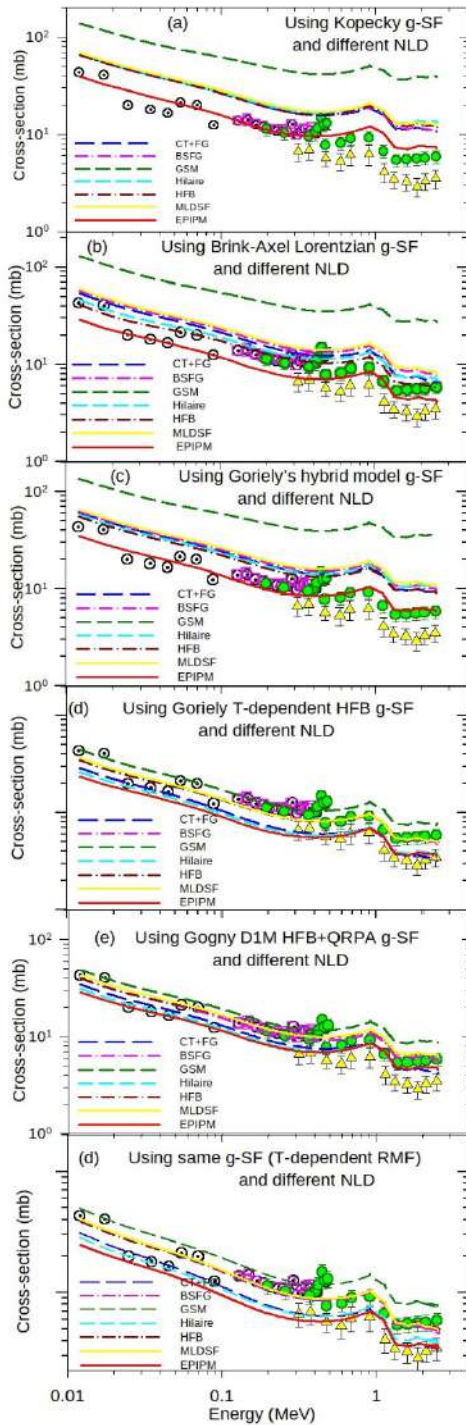


Fig. 1 Experimental cross-section [9] (different symbols denote data taken from different references in EXFOR) of $^{68}\text{Zn}(n,\gamma)^{69}\text{Zn}$ along with the results obtained from TALYS-1.95 calculations using different combinations of existing NLD and γ -SF and also along with the result by using the EP+IPM NLD and different existing γ -SFs (continuous line).

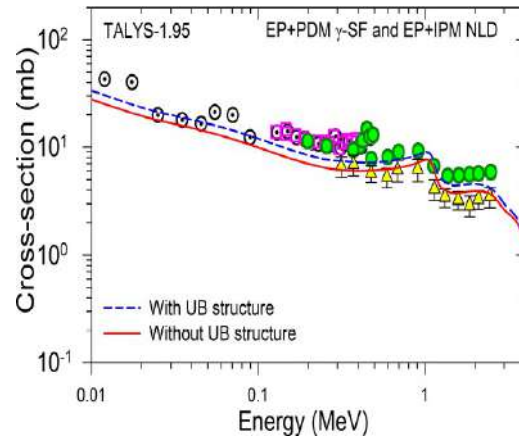


Fig. 2 Experimental cross-section of $^{68}\text{Zn}(n,\gamma)^{69}\text{Zn}$ along with the results obtained from TALYS-1.95 calculations using EP+IPM NLD and EP+PDM γ -SF with (dashed line) and without (continuous line) UB structures.

Acknowledgments

The authors acknowledge the financial support of the Science and Engineering Research Board (SERB), Government of India via Grant No. SUR/2022/000467 and Department of Science and Technology (DST)-WISE fellowship, Government of India via Grant No. DST/WISE-PhD/PM/2023/11(G). The authors (N.N.A and N.Q.H) acknowledge the financial support of Vietnam National Foundation for Science and Technology Development (NAFOSTED) under grant number NCU02-2022.38.

References

- [1] L. Wolfenstein, Phys. Rev. 82 (1951) 690.
- [2] A. C. Larsen et al., AIP Conference Proceedings 1377 (2011) 239246.
- [3] Rajkumar Santra et al., Phys. Lett. B 806 (2020) 135487.
- [4] M. Yigit and M. E. Korkmaz, Modern Physics Letters A Vol. 33, No. 26, (2018) 1850155.
- [5] N. Quang Hung, N. Dinh Dang, L.T. Quynh Huong, Phys. Rev. Lett. 118 (2017) 022502.
- [6] Senapati E et al 2023 J. Phys. G: Nucl. Part. Phys. 50 075104
- [7] Goryachev A M and Zalesnyy G N 1982 Voprosy Teoreticheskoy i Yadernoy Fiziki 8 121
- [8] Enakshi Senapati et al 2024 J. Phys. G: Nucl. Part. Phys. 51 115104.
- [9] Nuclear Data Sheets Volume 120, June 2014, Pages 272-276, [<https://nndc.bnl.gov/exfor>].