# Analysis of Λ-binding energies in the relativistic and non-relativistic approach

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## Introduction

The  $\Lambda$ -hypernuclear systems have been analyzed by means of both relativistic and non-relativistic approaches. Relativistic approach takes into account the spin-orbit force occurring naturally in the theory [1] and also plays an important role in nuclear saturation phenomena [2]. Hence, in relativistic calculations, a good reproduction of the  $\Lambda$ -binding energies ( $B_{\Lambda}$ ) is expected. Here, we make a comparative study of the  $B_{\Lambda}$  in hypernuclei using our non-relativistic approach [3] and the relativistic approach followed by Koutroulos and Grypeos [4].

### **Formulation**

In our non-relativistic phenomenological approach, we have obtained a semi-empirical formula for  $B_{\Lambda}$  using the point nucleon (N) density  $\rho_N(r)$  as an average of point proton density  $\rho_p(r)$  and point neutron density  $\rho_n(r)$ :

$$\rho_{N}(r) = \frac{Z}{A_{c}}\rho_{p}(r) + \frac{N}{A_{c}}\rho_{n}(r). \tag{1}$$

The single-particle  $\Lambda$ -nucleus potential is obtained by folding zero-range  $\Lambda N$  potential with the point nucleon density of the core nucleus. Solving the eigenvalue equation for  $B_{\Lambda}$ , in the approximation  $e^{-R/a} \ll 1$ , leads to the following semi-empirical formula [3]:

$$B_{\Lambda} = D_{\Lambda} - \frac{\hbar^{2} \pi^{2}}{2\mu_{\Lambda \Lambda}} \{ C_{0}^{'} A_{c}^{-2/3} - C_{1}^{'} A_{c}^{-1} + \cdots \}, \quad (2)$$

where the parameters are defined in ref. [3].

In the relativistic approach [4], the average local  $\Lambda$ -nucleus potential is constructed by means of an attractive scalar relativistic single particle potential  $U_s(r)$  and a repulsive relativistic single particle potential  $U_v(r)$  which is the fourth component of a vector potential. Writing the eigenvalue equation, in a way analogous to that

of the non-relativistic case and solving for  $B_{\Lambda}$  (for heavy hypernuclei), the following expression [4] is obtained:

$$\begin{split} B_{\Lambda}^{(0)} &= \frac{\mu c^2}{\lambda} \left\{ 1 + \lambda D_+ (2\mu c^2)^{-1} \right\} \left\{ 1 - \left[ 1 + 2\lambda (\mu c^2)^{-1} \left( \frac{\hbar^2 \pi^2 \lambda}{2\mu R^2} - D_+ \right) \right] \right. \\ &\left. \left. \left( 1 + \lambda D_+ (2\mu c^2)^{-1} \right)^{-2} \right]^{1/2} \right\}, \end{split}$$
(3)

where the symbols are defined in ref. [4].

Retaining the first term in the expansion of arctanx in powers of x, an improved form of  $B_{\Lambda}$  is obtained [4], which is given as

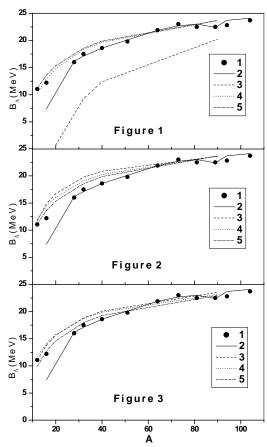
$$B_{\Lambda}^{(1)} = D_{+} - \frac{\hbar^{2}\pi^{2}}{2\mu g \left(1 + \left(\tilde{f}\eta_{0}R\right)^{-1}\right)^{2}R^{2}} \ , \tag{4} \label{eq:4}$$

where g,  $\tilde{f}$  and  $\eta_0$ , defined in ref. [4], depend upon  $B_\Lambda$  but their values are estimated by using an approximate expression  $B_{appr.} = D_+$  for  $B_\Lambda$ .

#### **Result and Discussion**

The radius and diffuseness parameters of the average nucleon density  $\rho_N(r)$  are obtained [3] from the least square fit to eq. (1) for nuclei over a large mass number range. With these parameters, the  $\chi^2$  fit to the ground state  $B_{\Lambda}$  values of  ${}^{28}_{\Lambda}Si$ ,  ${}^{32}_{\Lambda}S$ ,  ${}^{40}_{\Lambda}Ca$ ,  ${}^{51}_{\Lambda}V$  and  ${}^{89}_{\Lambda}Y$  is carried out [3] using eq. (2). The best fit value of  $D_{\Lambda}$  is 29.47 MeV. These parameters are then used to predict [3] the  $B_{\Lambda}$  values of  ${}^{16}_{\Lambda}O$  and the heavy and spallation hypernuclei corresponding to the mass number range A = 64, 73, 81, 94 and 104. The experimental  $B_{\Lambda}$  data of  ${}^{13}_{\Lambda}C$ ,  ${}^{16}_{\Lambda}O$ ,  ${}^{28}_{\Lambda}Si$ ,  ${}^{32}_{\Lambda}S$ ,  $^{40}_{\Lambda}\text{Ca}, \, ^{51}_{\Lambda}\text{V}, \, ^{89}_{\Lambda}\text{Y}$  and the upper limits of  $B_{\Lambda}$  in the case of the above mentioned mass number range, are shown as 1 in all the given figures. While our calculated  $B_{\Lambda}$  values obtained from fitting [3], along with the predicted values, are represented as 2 in all figures.

The  $B_{\Lambda}$  values calculated in ref. [4], are shown in the figures for  $A=12,\ 16,\ 20,\ 32,\ 40$  and 90. The calculated  $B_{\Lambda}$  values [4] for A=140 and 208 are excluded from the plots as their experimental values are unavailable. The  $B_{\Lambda}^{(0)}$ ,  $B_{\Lambda}^{(1)}$  and  $B_{\Lambda}$  (exact) values, in ref.[4], with  $D_{-}=443$  MeV,  $D_{+}=30.77$  MeV and  $r_{0}=1.022$  fm, are plotted as 3, 4 and 5 in Fig. 1. These values of  $B_{\Lambda}$ 's, further calculated in ref. [4], using  $D_{-}=443$  MeV and the corresponding best fit values of  $D_{+}$  and  $r_{0}$ , are plotted in Fig. 2, as 3, 4 and 5.



In Fig. 3, the best fit values [4] of  $B_{\Lambda}^{(1)}$  for  $D_{-}=412.84$  MeV,  $D_{+}=29.57$  MeV and  $r_{0}=1.132$  fm are plotted as 3. The  $B_{\Lambda}^{(1)}$  values [4] with  $D_{-}=443$  (fixed) MeV,  $D_{+}=29.03$  MeV,  $r_{0}=1.147$  fm and  $m^{*}=0.788$  m are plotted as 4. The calculated values [4] of  $B_{\Lambda}^{(1)}$ , with the parameters  $D_{-}=590.15$  MeV,  $D_{+}=29.50$  MeV,  $r_{0}=1.123$  fm,  $m^{*}=0.722$ m, obtained by least

square fitting of experimental data, are plotted as 5 in Fig. 3.

From Fig. 1, we can see that the  $B_{\Lambda}^{(0)}$  values (shown as 3), are way-off from the experimental data, while the  $B_{\Lambda}^{(1)}$  and  $B_{\Lambda}$  (exact) values (shown as 4 and 5), are quite close to each other but differ slightly from the experimental data for comparatively lower mass numbers. The  $B_{\Lambda}$  values obtained by us [3] (shown as 2 in all the figures) give a fairly good account of the experimental data over a wide range of mass numbers. The difference in the predicted  $B_{\Lambda}$  of  $^{16}_{\Lambda}O$  is not surprising as our semi-empirical formula for  $B_{\Lambda}$  is valid for heavy hypernuclei. The best fit values of  $B_{\Lambda}^{(0)}$ ,  $B_{\Lambda}^{(1)}$  as well as  $B_{\Lambda}$  (exact) in Fig. 2, (represented as 3, 4 and 5), are comparatively not so good for A< 60.

The  $B_{\Lambda}^{(0)}$  values (shown as 3) are quite unrealistic in Fig. 1 and show considerable deviation from the experimental data in Fig. 2. The  $B_{\Lambda}^{(1)}$  values (shown as 5), in Fig. 3, are comparatively better, while the others (shown as 3 and 4), are more or less same but show deviations from the experimental data for lower mass number range. The D<sub>2</sub> parameter seems to play a significant role in the fitting. With the higher value of D<sub>2</sub> and only marginal changes in other parameters, the  $B_{\Lambda}$  values are reproduced fairly well (shown as 5) in Fig. 3.

In comparison to the relativistic case [4] our non-relativistic results [3] give a much better reproduction of the experimental data, as is evident from Figs. 1, 2 and 3. This anomaly might be due to the approximations chosen in relativistic semi-empirical mass formula [4] for  $B_{\Lambda}$  of heavy hypernuclei. However, more information is needed to draw any definite conclusion about the significance of relativistic approach for the determination of  $B_{\Lambda}$ .

## References

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