

Studies on probable chances of decay of superheavy element ³⁰⁴120

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Introduction

Theoretical studies on superheavy nuclei (SHN) have become a prime concern in recent years so that we can develop stable nuclei in the superheavy region and have a better understanding of the nuclear structure. SHN up to Z=118 has been experimentally synthesized so far using the cold and hot fusion evaporation technique and attempt to study Z=120 have been reported. Superheavy element ³⁰⁴120 is an isotope of great interest among researchers since it has magic number of proton (Z=120) and neutron (N=184).

In the present work, we aim to study the possibilities of SHN ³⁰⁴120 to emit alpha and heavier cluster using the Modified Generalized Liquid drop model (MGLDM) [1] with Q value dependent pre-formation factor [2] and also we calculate its decay modes by comparing alpha decay half-lives with spontaneous fission half-lives computed using shell effect and mass inertia dependent formula [3, 4].

Theory

In MGLDM, for a deformed nucleus, the macroscopic energy is defined as,

$$E = E_V + E_S + E_C + E_R + E_P \quad (1)$$

Here the terms E_V , E_S , E_C , E_R and E_P represents the volume, surface, Coulomb, rotational and nuclear proximity energy terms respectively.

The barrier penetrability P is calculated with the action integral

$$P = \exp \left\{ -\frac{2}{\hbar} \int_{R_{in}}^{R_{out}} \sqrt{2B(r)[E(r) - E(sphere)]} dr \right\}$$

Where $R_{in} = R_1 + R_2$, $B(r) = \mu$, the reduced mass and $R_{out} = e^2 Z_1 Z_2 / Q$. R_1 , R_2 are the

radius of the daughter nuclei and emitted cluster respectively, Q the released energy.

The partial half-life is related to the decay constant λ by

$$T_{1/2} = \left(\frac{\ln 2}{\lambda} \right) = \left(\frac{\ln 2}{\nu P_C P} \right) \quad (2)$$

The assault frequency ν has been taken as 10^{20} s^{-1} and the pre-formation factor is given as

$$P_C = 10^{aQ+bQ^2+c} \quad (3)$$

With $a = -0.25736$, $b = 6.37291 \times 10^{-4}$, $c = 3.35106$ and Q is the Q value of the reaction.

Results and discussion

We have calculated all possible cluster daughter combination possible in the decay of ³⁰⁴120 using the concept of cold valley reaction plot. Half-lives and branching ratio of all possible fragment combinations of SHN ³⁰⁴120 are calculated. Half-lives are computed using MGLDM with Q value dependent pre-formation factor. We considered only those splitting of ³⁰⁴120 which are within the experimental half-life limit (less than 10^{30} s) and branching ratio limit (down to 10^{-19}). Among all possible cluster emission, ¹²¹Cd with ¹⁸³Hf daughter nuclei is found to be the most probable heavy cluster decay of ³⁰⁴120 with half-lives comparable with alpha decay half-lives. Half-lives of some important splitting of ³⁰⁴120 are shown in table 1. Also in all possible fragment combinations, ¹³³Te with ¹⁷¹Er daughter nuclei, ¹³⁵I with ¹⁶⁹Ho daughter nuclei and ¹³⁸Xe cluster with ¹⁶⁶Dy daughter nuclei are the most favorable heavy cluster reaction possible, with minimum half-life. These heavy clusters have neutron number N= 82 or near to it, which is a magic number, thereby proving the role of neutron shell closure in heavy cluster decay.

A graph with cluster size along X axis and logarithm of half-lives using MGLDM with Q dependent preformation factor along Y axis, of splitting of $^{304}120$ is plotted in figure 1.

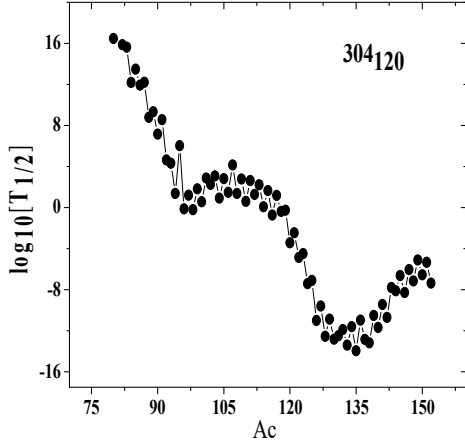


Figure 1. Graph connecting cluster size vs. logarithm of half-life of splitting of $^{304}120$.

Table 1: Splitting of $^{304}120$ whose half-lives are comparable with alpha decay half-lives and those with minimum half-life are given.

Daughter Nuclei	Probable Cluster	Q Value (MeV)	$T_{1/2}$ (s)
^{300}Og	^4He	11.52	0.006648
^{121}Cd	^{183}Hf	344.62	0.003418
^{133}Te	^{171}Er	360.93	3.86E-14
^{135}I	^{169}Ho	362.85	1.13E-14
^{138}Xe	^{166}Dy	362.83	6.50E-14

We also calculated the decay modes of $^{304}120$ by comparing alpha decay half-life with spontaneous fission half-lives and are listed in table 2. Alpha decay half-lives are calculated using the MGLDM and Spontaneous fission half-lives are calculated using the shell effect and mass inertia dependent equation [3, 4] and is,

$$\log_{10} [T_{1/2} (yr)] = c_1 + c_2 \left(\frac{z^2}{(1 - kI^2)A} \right) + c_3 \left(\frac{z^2}{(1 - kI^2)A} \right)^2 + c_4 E_{sh} + c_5 I_{rigid} + h_i$$

The rigid body mass inertia of the nucleus [5, 6] is given as

$$I_{rigid} = B_{rigid} [1 + 0.31\beta_2 + 0.44\beta_2^2 + \dots] \quad (4)$$

With the mass parameter,

$$B_{rigid} = \frac{2}{5} MR^2 = 0.0138 A^{5/3} (\hbar^2 / \text{MeV}) \quad (5)$$

The value of constants are $c_1 = 1208.763104$, $c_2 = -49.26439288$, $c_3 = 0.486222575$, $c_4 = 3.557962857$, $c_5 = 0.04292571494$, $k = 2.6$ and h_i , is blocking effect given in Ref. [3, 4].

From table it is clear that for first 4 decays, alpha decay half-life is less compared to spontaneous fission half-life and in the fifth case, spontaneous fission half-life is less. Thus without any doubt, we can say that for superheavy element $^{304}120$, decays by 4 alpha chains followed by spontaneous fission.

Table 2: Comparison of alpha decay half-lives with spontaneous fission half-lives and the decay modes of $^{304}120$.

Parent Nuclei	Q_α (MeV)	T_{SF} (s)	T_α (s)	Mode of decay
$^{304}120$	11.52	48.85028	0.00665	α
^{300}Og	11.04	7.56E+03	0.02792	α
^{296}Lv	10.51	2.13E+06	0.17724	α
^{292}Fl	8.38	7.99E+06	2.83E+05	α
^{288}Cn	7.95	2.19537	2.20E+04	SF

References

- [1] K. P. Santhosh, C. Nithya, H. Hassanabadi and Dashty T. Akrawy, Phys. Rev. C 98, (2018): 024625.
- [2] K. P. Santhosh and Tinu Ann Jose, Phys. Rev. C 99, (2019): 064604.
- [3] K. P. Santhosh, Tinu Ann Jose and N. K. Deepak, Phys. Rev. C 103, (2021): 064612
- [4] K. P. Santhosh, C. Nithya and Tinu Ann Jose, Phys. Rev. C 104, (2021): 024617.
- [5] A. Bohr and B. R. Mottelson, Dan. Mat. Fys. Medd. 30, (1955): 1.
- [6] J. M. Allmond and J. L. Wood, Phys. Lett. B 767, (2017): 226.